

# Magnetic confinement of charged particles

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# Overview

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- Motivation for studying single particle orbits in magnetic fields
- Motion in a straight uniform B-field (basic review)
- The guiding center approximation
- Motion including a static electric field, or gravity
- Motion in inhomogeneous and bent B-fields and the first adiabatic invariant
- Mirror confinement
- Time variation (no derivation)

# The need for magnetic field confinement

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- The easiest fusion process to reach is D-T fusion
- This requires particle kinetic energies in the range 10-100 keV
- Even at the particle energy of peak D-T fusion reactivity, non-fusion collisions (scattering) dominate over the fusion collisions by two orders of magnitude
  - Must confine plasma at  $T > 10$  keV ( $\sim 120$  M Kelvin) for many collisions
- Thermal speed of D and T is on the order of  $10^6$  m/s at these temperatures (and even higher for electrons) –  $\mu$ s confinement if you have no confining field?
- Electric fields alone won't work: Confine only one species
- Magnetic fields may work (must be bent!)
- Gravity works on the sun, but not on Earth

# Charged particle motion in a straight magnetic field

- A charged particle performs a screw-like path if it is confined by a straight uniform magnetic field and it feels no other forces
- Start with Newton's 2<sup>nd</sup> law and the Lorentz force:

$$\vec{F} = m\vec{a} = q\vec{v} \times \vec{B}$$

$$\vec{B} = B_0 \hat{z}$$



# Charged particle motion in a straight magnetic field

- Newton's 2<sup>nd</sup> law written coordinate by coordinate:

$$\vec{F} = m\vec{a} = q\vec{v} \times \vec{B} \Rightarrow$$

$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} v_x$$

$$\frac{dv_z}{dt} = 0$$

# Charged particle motion in a straight magnetic field

- Newton's 2<sup>nd</sup> law written coordinate by coordinate:
- $qB_0/m$  is an inverse time scale

$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} v_x$$

$$\frac{dv_z}{dt} = 0$$

# Charged particle motion in a straight magnetic field

- $qB_0/m$  is an inverse time scale: give it it's own symbol

$$\frac{dv_x}{dt} = \omega_c v_y$$

$$\frac{dv_y}{dt} = -\omega_c v_x$$

$$\frac{dv_z}{dt} = 0$$

$$\omega_c = \frac{qB_0}{m} \quad - \quad \text{sometimes } \omega_c = \frac{|q|B_0}{m}$$

# Charged particle motion in a straight magnetic field

- Decouple  $v_x$  and  $v_y$  equations:

$$\frac{d}{dt} \left[ \frac{dv_x}{dt} = \omega_c v_y \right]$$

$$\frac{dv_y}{dt} = -\omega_c v_x$$

$$\frac{dv_z}{dt} = 0$$

$$\omega_c = \frac{qB_0}{m} \quad - \quad \text{sometimes } \omega_c = \frac{|q|B_0}{m}$$



# Charged particle motion in a straight magnetic field

- Eliminate  $v_y$  from  $v_x$  equation by differentiation and substitution
- $V_z$  equation is trivial

$$\left. \begin{aligned} \frac{d^2 v_x}{dt^2} &= \omega_c \frac{dv_y}{dt} \\ \frac{dv_y}{dt} &= -\omega_c v_x \end{aligned} \right\} \Rightarrow \frac{d^2 v_x}{dt^2} = -\omega_c^2 v_x$$

$$\frac{dv_z}{dt} = 0 \Rightarrow v_z = \text{constant} = v_{\parallel}$$

$$\omega_c^2 = \left( \frac{qB_0}{m} \right)^2$$

# Charged particle motion in a straight magnetic field

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- We recognize the simple harmonic oscillator for  $v_x$
- Find  $v_y$  by differentiation
- $v_z$  equation is trivial:

$$\frac{dv_z}{dt} = 0 \Rightarrow v_z = \text{constant} = v_{\parallel}$$

$$\omega_c^2 = \left( \frac{qB_0}{m} \right)^2$$

Lecture on guiding center approximation

# Charged particle motion in a straight magnetic field

- Eliminate  $v_y$  from  $v_x$  equation by differentiation and substitution:

$$\left. \begin{aligned} \frac{dv_x}{dt} &= \omega_c v_y \\ \frac{dv_y}{dt} &= -\omega_c v_x \end{aligned} \right\} \Rightarrow \begin{aligned} \frac{d^2 v_x}{dt^2} &= -\omega_c^2 v_x \\ v_y &= \frac{1}{\omega_c} \frac{dv_x}{dt} \end{aligned}$$

- We recognize the simple harmonic oscillator for  $v_x$
- Find  $v_y$  by differentiation
- $v_z$  equation is trivial:

$$\frac{dv_z}{dt} = 0 \Rightarrow v_z = \text{constant} = v_{\parallel}$$

$$\omega_c^2 = \left( \frac{qB_0}{m} \right)^2$$

# Charged particle motion in a straight magnetic field

- Velocity components:

$$v_x = v_{\perp} \cos(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta)$$

$$v_z = v_{\parallel}$$

$$\omega_c^2 = \left(\frac{qB_0}{m}\right)^2$$

- Next step: integrate to get position

# Charged particle motion in a straight magnetic field

- Integrate in time to get position
- Define Larmor radius and guiding center:

$$v_x = v_{\perp} \cos(\omega_c t + \delta) \Rightarrow x = x_{gc} + \frac{v_{\perp}}{\omega_c} \sin(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta) \Rightarrow y = y_{gc} + \frac{v_{\perp}}{\omega_c} \cos(\omega_c t + \delta)$$

$$v_z = v_{\parallel} \Rightarrow z = z_0 + v_{\parallel} t$$

$$\omega_c = \frac{qB_0}{m} \text{ (cyclotron frequency)}, f_c = \omega_c / 2\pi$$

$$r_L = \left| \frac{v_{\perp}}{\omega_c} \right| = \left| \frac{mv_{\perp}}{qB_0} \right| \text{ (Larmor radius, gyroradius)}$$

# The guiding center approximation:

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- If the Larmor radius is on the order of the size of the confining field, then only collisionless orbits are confined:
  - After a collision, the particle will have a new guiding center, about one Larmor radius away from the original guiding center
  - If this new Larmor orbit intersects material walls or extends to regions of much lower B-field, the particle is not confined
- In order to confine charged particles magnetically for many collision times, the Larmor radius must be small!
- When the Larmor radius is small, and the cyclotron frequency is large:
- one can derive analytic formulas for the time evolution of the guiding center  $(x_{gc}, y_{gc}, z)$ , averaging over the gyration

$$r_L \ll B / \nabla B \text{ and } \omega_c \gg \frac{\partial B}{\partial t} \frac{1}{B} (*)$$

# The guiding center approximation is often necessary:

- For a 1 eV electron in a  $B=1$  T field, the cyclotron frequency is large and the Larmor radius small:

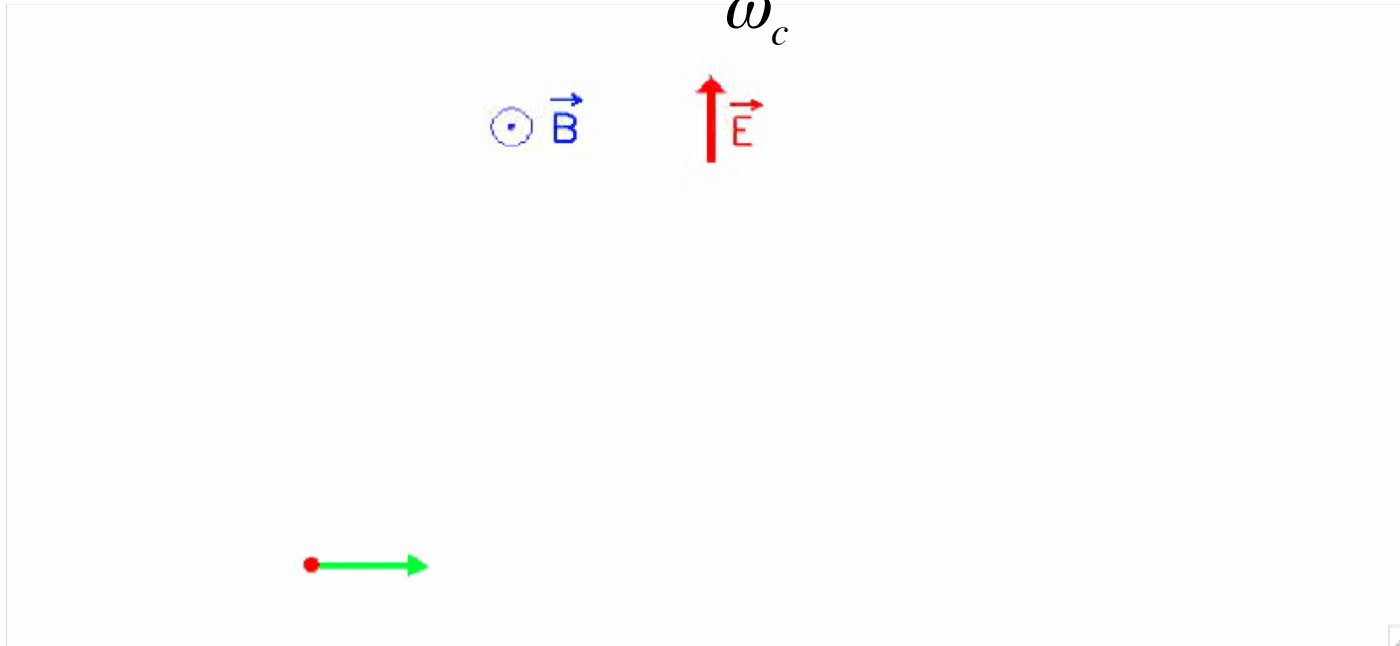
$$\omega_c \approx 1.8 \times 10^{11} \text{ s}^{-1}$$

$$r_L \approx 3 \times 10^{-6} \text{ m}$$

- Just to follow the electron for one microsecond requires  $>10^6$  time steps if a simple numerical scheme is used.
- Almost the full computational effort is spent calculating the circular motion....
- Averaging over the gyromotion allows fast and accurate calculations of the motion of charged particles in a magnetic field, both analytic and numerical
- Essentially, the gyrating particle is replaced by a charged ( $q$ ), massive ( $m$ ) ring of current ( $I=e\omega_c/2\pi$ ), with its center at the particle's gyrocenter.
- We will do this for a few important cases in the following:

# ExB drift

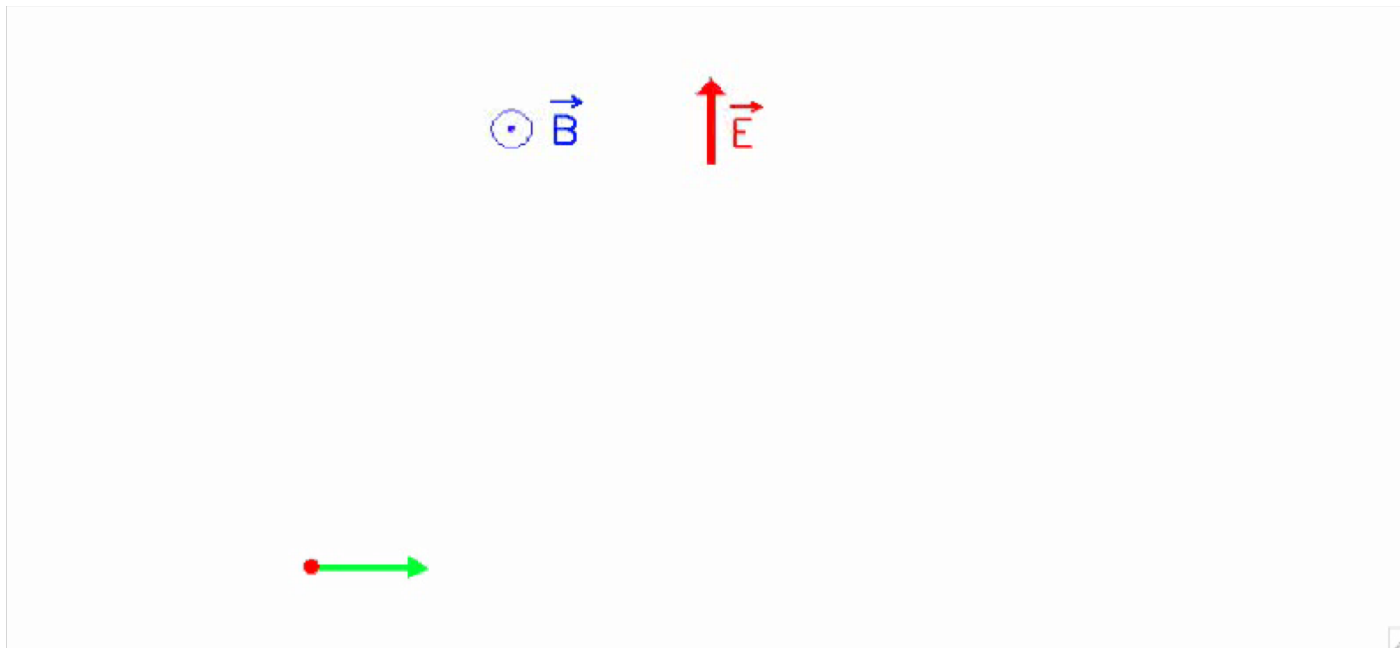
- Since particles are charged, the electric field naturally enters the equations (collective phenomena, external confining fields, single particle Coulomb interactions)
- Electric field component along B gives simple acceleration or deceleration
- Electric field component perpendicular to B is more interesting:
- Example below: Electron  $r_L = \frac{v_{\perp}}{\omega_c}$





# ExB drift

- Net effect is a motion in the  $\mathbf{E} \times \mathbf{B}$  direction which has a steady state component
- Let's calculate this drift:



# ExB drift

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- Newton's 2<sup>nd</sup> law written coordinate by coordinate (again)
- Let's add an electric field now

$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} v_x$$

$$\frac{dv_z}{dt} = 0$$

# ExB drift

---

$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} v_x + \frac{q}{m} E_y$$

$$\frac{dv_z}{dt} = 0$$

# ExB drift

---

- $E_y/B_0$  is a velocity,  $v_E$ ; it is constant since we assumed E and B constant.

$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} \left( v_x - \frac{E_y}{B_0} \right)$$

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## ExB drift

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- $E/B$  is a velocity,  $v_E$ ; it is constant since we assumed  $E$  and  $B$  constant.
- Since  $v_E$  is constant, we can subtract it inside the  $d/dt$  of  $x$ -equation

$$\frac{d(v_x - v_E)}{dt} = \frac{qB_0}{m} v_y$$

$$\frac{dv_y}{dt} = -\frac{qB_0}{m} (v_x - v_E)$$

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**Now**

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$$\frac{dv_x}{dt} = \frac{qB_0}{m} v_y$$

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**Before**



## ExB drift

---

- It's clear we have the same equation as before, just replacing  $v_x$  by  $v_x - v_E$

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**Before**

# ExB drift

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- It's clear we have the same equation as before, just replacing  $v_x$  by  $v_x - v_E$

$$v_x - v_E = v_{\perp} \cos(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta)$$

$$v_z = v_{\parallel}$$

$$\omega_c^2 = \left(\frac{qB_0}{m}\right)^2$$

**Now**

$$v_x = v_{\perp} \cos(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta)$$

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**Before**

# ExB drift

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- It's clear we have the same equation as before, just replacing  $v_x$  by  $v_x - v_E$

$$v_x = v_E + v_{\perp} \cos(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta)$$

$$v_z = v_{\parallel}$$

$$\omega_c^2 = \left(\frac{qB_0}{m}\right)^2$$

**Now**

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**Before**

# ExB drift

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$$v_x = \frac{E_y}{B_0} + v_{\perp} \cos(\omega_c t + \delta)$$

$$v_y = -v_{\perp} \sin(\omega_c t + \delta)$$

$$v_z = v_{\parallel}$$

$$\omega_c^2 = \left(\frac{qB_0}{m}\right)^2$$

**Now**

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$$\omega_c^2 = \left(\frac{qB_0}{m}\right)^2$$

**Before**

## ExB drift: Coordinate-free formulation

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- For situations with constant uniform  $E$  and  $B$  fields, we can always define a local coordinate system where  $z$  is in the  $B$ -field direction and  $y$  is in the direction of the component of  $E$  perpendicular to  $B$ ; hence, our derivation is valid in any coordinate system. The coordinate free formula for  $v_E$  is:

$$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2} = -\frac{\nabla\phi \times \vec{B}}{B^2}$$

- The drift is independent of the particle! No reference to  $q$  or  $m$ 
  - Same for ions and electrons
- The drift goes along constant  $\phi$  surfaces – does not change the electrostatic energy of the particle

# ExB drift: Coordinate-free derivation

- We can also prove that the particles ExB drift without using coordinates

$$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2} = -\frac{\nabla\phi \times \vec{B}}{B^2}$$

- We know what we are looking for:  $\vec{v}_\perp = \vec{v}_g + \vec{v}_E$

$$\begin{aligned} m \frac{d\vec{v}}{dt} &= q(\vec{E}_\perp + \vec{v} \times \vec{B}) \Rightarrow m \frac{d\vec{v}_\perp}{dt} = q(\vec{E}_\perp + \vec{v}_E \times \vec{B} + \vec{v}_g) \\ &= q\left(\vec{E}_\perp + (\vec{E} \times \vec{B}) \times \frac{\vec{B}}{B^2} + \vec{v}_g\right) = q\vec{v}_g \times \vec{B} \Leftrightarrow m \frac{d\vec{v}_g}{dt} = q\vec{v}_g \times \vec{B} \end{aligned}$$

## ExB drift

---

$$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2} = -\frac{\nabla\phi \times \vec{B}}{B^2}$$

- The drift is independent of the particle! There is no reference to  $q$  or  $m$
- Same for ions and electrons, pions, all charged particles....why?

Answer: In the inertial frame that moves at the ExB velocity, there is no E-field!

$$E' = \gamma(\vec{E} + \vec{v} \times \vec{B}) + (1 - \gamma) \frac{\vec{v} \cdot \vec{E}}{v^2} \vec{v} =$$

Lorentz transform:

$$\gamma(\vec{E} + \frac{\vec{E} \times \vec{B}}{B^2} \times \vec{B}) = \gamma(\vec{E} - \vec{E}) = 0$$

- A charged particle therefore performs simple cyclotron motion in that frame (as long as  $v_E = E/B < c$ )
- Exercise: What happens when  $E/B > c$ ?



# FxB drift

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- The derivation we did only used Newton's 2<sup>nd</sup> law – no reference to the Lorentz transform or Maxwell's equations
  - (then afterwards it was realized that we could have used Lorentz)
- But this is actually an advantage: Our derivation can be trivially extended to any other constant perpendicular force acting on our particle:

$$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2}$$

$$\vec{v}_F = \frac{\vec{F} \times \vec{B}}{qB^2}$$

- The general force drift cares about the particle's charge, as one would expect. Example  $F=mg$  leads to a gravitational drift in opposite directions for electrons and ions.

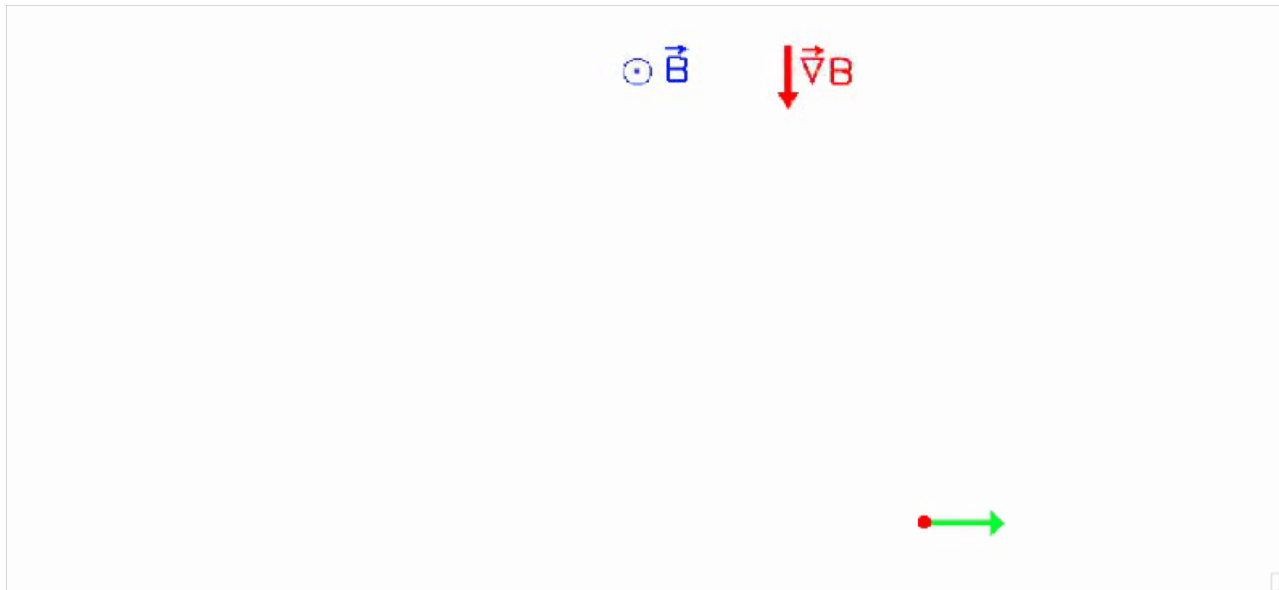
# Non-uniform B-field

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- In some traps, the magnetic field is non-uniform
- We assumed  $B$  straight and uniform
- What happens when (for example) the B-field strength changes spatially?

# Non-uniform B-field

- In some traps, the magnetic field is non-uniform
- We assumed B straight and uniform
- What happens when (for example) the B-field strength changes spatially?
- Assume for simplicity here: B-field is straight but increases in strength

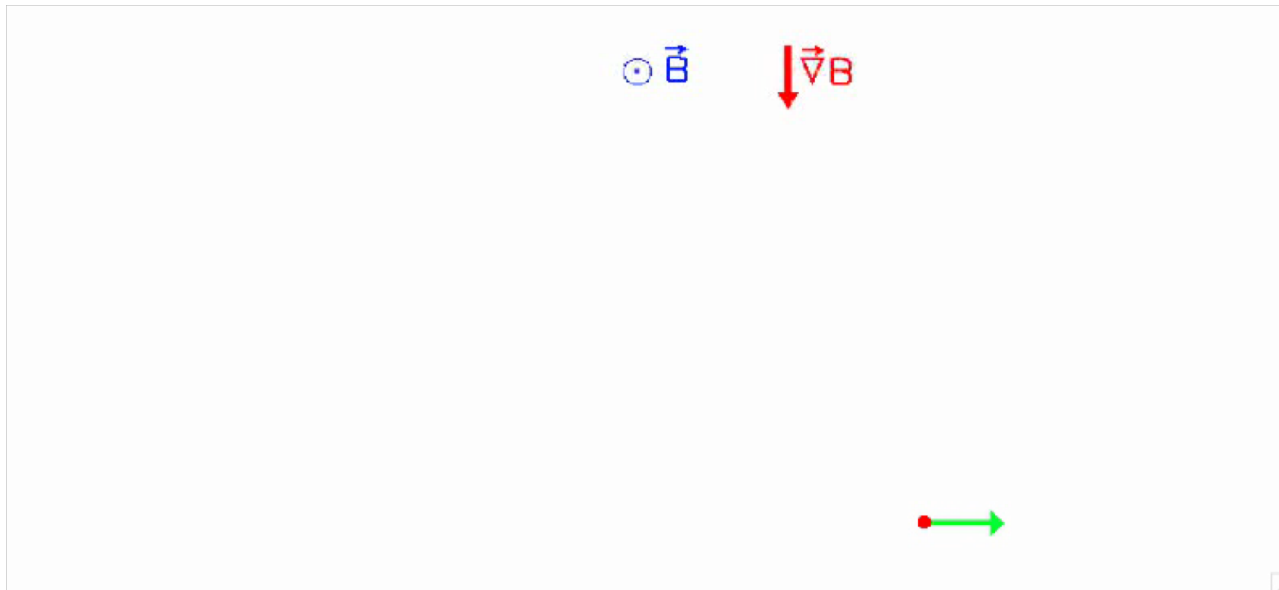


# Non-uniform B-field

- One can derive this drift by Taylor expanding the B-field, taking advantage of the smallness of the Larmor radius (keeping only 1<sup>st</sup> order terms)

$$B_z(x, y, z) = B_0 + \frac{\partial B_z}{\partial y} (y - y_{gc}) + O(\varepsilon^2)$$

$$\varepsilon = \left| \frac{\nabla B}{B} \right| r_L \ll 1$$



# Adiabatic invariants

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- Instead of Taylor expanding, it is also possible to derive this drift much faster by introducing the first adiabatic invariant  $\mu$ :
  - This invariant is very useful in several contexts
- Background:
  - The concept of adiabatic invariants is known from analytic mechanics
  - Assume a particle performs periodic motion in one coordinate  $q$
  - Then one can define the action as:  $\oint p_q dq$ 
    - Here  $p_q$  is the generalized momentum associated with  $q$
  - If one perturbs the periodic motion by a small amount  $\varepsilon$ , the action remains conserved, to all powers in  $\varepsilon$
- We have already one periodic motion – the gyration. The coordinate for the gyration is  $\theta$ , and  $p_\theta = mv_\theta r$  is the associated generalized momentum (we recognize it's just the angular momentum in the gyration)

# The first adiabatic invariant

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$$\oint p_q dq = \int_0^{2\pi} m v_\theta r d\theta = \int_0^{2\pi} m v_\perp r_L d\theta = 2\pi m v_\perp r_L =$$

$$2\pi m v_\perp \frac{m v_\perp}{q B} = \frac{4\pi m}{q} \frac{\frac{1}{2} m v_\perp^2}{B} = \text{constant}$$

$$\text{so } \mu = \frac{\frac{1}{2} m v_\perp^2}{B} = \text{constant}$$

- $\mu$  is conserved – it is actually the magnetic dipole moment of the charged particle, if we consider the particle as a charged current ring with radius  $r_L$

$$IA = \frac{q\omega_c}{2\pi} \pi r_L^2 = \frac{q^2 B}{2\pi m} \pi \left(\frac{m v_\perp}{q B}\right)^2 = \frac{m v_\perp^2}{2B} = \mu$$

# The first adiabatic invariant is the dipole moment

$$IA = \frac{q\omega_c}{2\pi} \pi r_L^2 = \frac{q^2 B}{2\pi m} \pi \left( \frac{mv_\perp}{qB} \right)^2 = \frac{mv_\perp^2}{2B} = \mu$$

- This magnetic dipole is anti-aligned with the magnetic field (a plasma is diamagnetic)
- A magnetic dipole with strength  $\mu$  embedded in a magnetic field  $B$  anti-aligned to the dipole has potential energy  $\mu B$ , so it feels a force

This is the so-called mirror force  $\vec{F} = -\mu \nabla B$

This force also works for neutral particles as long as they have a magnetic dipole moment (Example: Antihydrogen)

If the force is along  $B$ , it can provide some confinement along the field lines for charged particles – they can be reflected by a magnetic mirror

(Mirror confinement was attempted for fusion but is not pursued much these days)

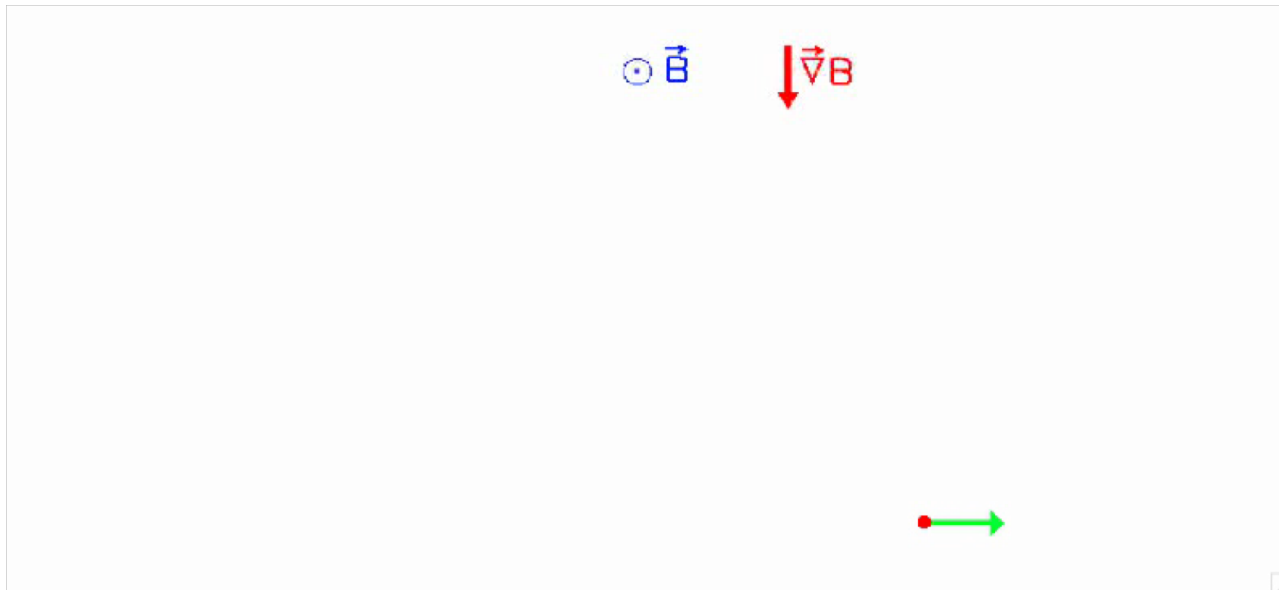
If the force is perpendicular to  $B$ , then we get a drift

# Non-uniform B-field strength

- We use now the FxB formula we derived earlier:

$$\vec{F} = -\mu \nabla B = -\mu \frac{dB_z}{dy} \hat{y}$$

$$\vec{v}_{\nabla B} = \frac{\vec{F} \times \vec{B}}{qB^2} = \frac{-\mu \nabla B \times \vec{B}}{qB^2} = \frac{mv_{\perp}^2}{2B} \frac{\vec{B} \times \nabla B}{qB^2}$$





# Non-uniform B-field direction: Curvature drift

- If the magnetic field strength is inhomogeneous, the magnetic field is usually also curved
  - It has to be curved if it's inhomogeneous, unless you have significant currents
- In the guiding center approximation, the zeroth order motion is along the magnetic field.
- So if the magnetic field is curved, the particle feels a centrifugal force:

$$\vec{F}_C = \frac{mv_{\parallel}^2}{R_C} \frac{\vec{R}_C}{R_C}$$
$$\vec{v}_{R_C} = \frac{\vec{F}_C \times \vec{B}}{qB^2} = \frac{mv_{\parallel}^2}{R_C^2} \frac{\vec{R}_C \times \vec{B}}{qB^2}$$

# Combining the two drifts

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- With a bit of algebra, we can combine the grad B and curvature drifts into one formula – assuming that the current density is negligible. This is not universally true but often enough that it is useful to derive this combined formula.

$$\begin{aligned}\nabla \times \vec{B} &= \mu_0 \vec{j} = 0 \\ \nabla \cdot \vec{B} &= 0\end{aligned}$$

- If the magnetic field has curvature, we can go into a local cylindrical coordinate system with the axis given by the axis for the radius of curvature (blackboard). In that coordinate system:

$$\vec{B} = B_r(r, \theta, z) \hat{r} + B_\theta(r, \theta, z) \hat{\theta} + B_z(r, \theta, z) \hat{z} = B_\theta(r, \theta, z) \hat{\theta} = B_\theta(r, \theta) \hat{\theta}$$

# Combining the two drifts

$$\vec{B} = B_r(r, \theta, z)\hat{r} + B_\theta(r, \theta, z)\hat{\theta} + B_z(r, \theta, z)\hat{z} = B_\theta(r, \theta, z)\hat{\theta} = B_\theta(r, \theta)\hat{\theta}$$

- We need Goldston and Rutherford (or another formula book) to write the differential operators in cylindrical coordinates:

$$\nabla \cdot \vec{B} = 0 \Rightarrow \frac{1}{r} \frac{\partial(rB_r)}{\partial r} + \frac{1}{r} \frac{\partial B_\theta}{\partial \theta} + \frac{\partial B_z}{\partial z} = 0 \Rightarrow \frac{\partial B_\theta}{\partial \theta} = 0 \Rightarrow \vec{B} = B_\theta(r)\hat{\theta}$$

- This helps us significantly simplify the curl equation, which is otherwise terribly complicated:

$$\nabla \times \vec{B} = \frac{1}{r} \left( \frac{\partial B_z}{\partial \theta} - \frac{\partial B_\theta}{\partial z} \right) \hat{r} + \left( \frac{\partial B_r}{\partial z} - \frac{\partial B_z}{\partial r} \right) \hat{\theta} + \frac{1}{r} \left( \frac{\partial(rB_\theta)}{\partial r} - \frac{\partial B_r}{\partial \theta} \right) \hat{z}$$

$$\nabla \times \vec{B} = \frac{1}{r} \left( \frac{\partial(rB_\theta)}{\partial r} \right) \hat{z} = 0 \Leftrightarrow rB_\theta = \text{constant (c)}$$

$$\Leftrightarrow B_\theta = \frac{c}{r}$$

# Combining the two drifts

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$$B = B_\theta = \frac{c}{r} \Rightarrow \nabla B = -\frac{c}{r^2} r$$

$$\frac{\nabla B}{B} = \frac{-\frac{c}{r^2} \hat{r}}{\frac{c}{r}} = \frac{-\hat{r}}{r} = \frac{-\vec{R}_c}{R_c^2}$$

# Combining the two drifts

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- Thus, we have shown in the last few slides that:

$$\left. \begin{array}{l} \nabla \times \vec{B} = \mu_0 \vec{j} = 0 \\ \nabla \cdot \vec{B} = 0 \end{array} \right\} \Rightarrow \frac{\nabla B}{B} = -\frac{\vec{R}_C}{R_C^2}$$

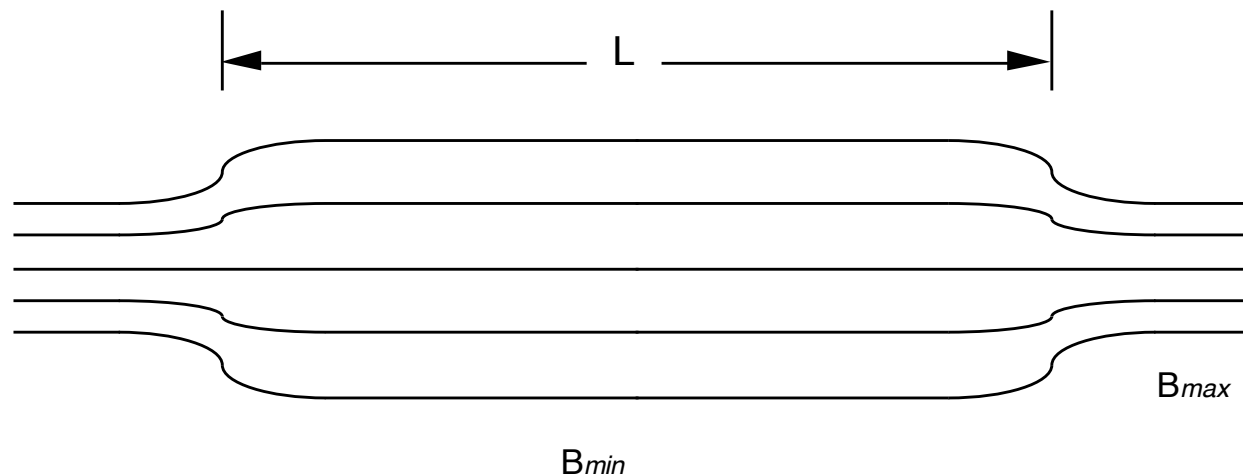
- Therefore we can combine the two magnetic non-uniformity drifts into one formula:

$$\vec{v}_{\nabla B + R_C} = \frac{2mv_{\parallel}^2 + mv_{\perp}^2}{2B} \frac{\vec{B} \times \nabla B}{qB^2}$$

- These two drifts add up – no magic cancellation
- The drifts are “slow” in general, verifying our perturbative approach:
  - E=100 eV electron in B=1 T and  $R_C = 1$  m:
    - $v_{R_C} \sim 100$  m/s vs  $\sim 5 \cdot 10^5$  m/s free streaming velocity and  $r_L \sim 3 \mu\text{m}$

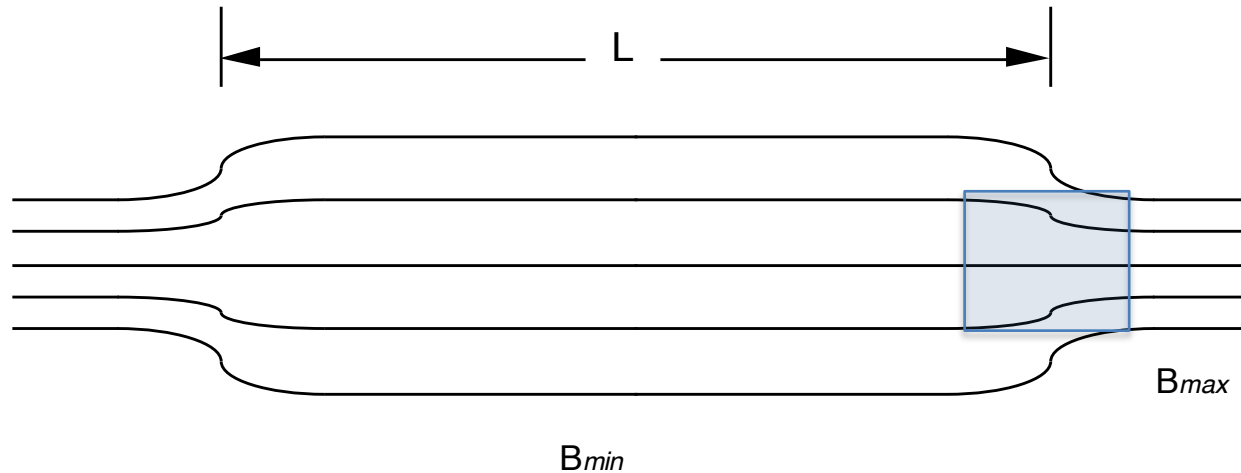
# Mirror force – mirror confinement

- The mirror force  $\mu \nabla B$  for a guiding center particle can also be parallel to  $B$ , and thereby provide confinement of guiding center particles in the third direction
- This immediately brings up the question whether this is physical – how can  $\vec{v} \times \vec{B}$  have a component parallel to  $\vec{B}$  ?



- Look at a particle whose guiding center is on the axis – the straight field line - of the magnetic configuration above

# Mirror force for a generic situation



- The particle is in a region of converging field lines:

$$B_z = B_z(z)$$

$$\nabla \cdot \vec{B} = 0$$

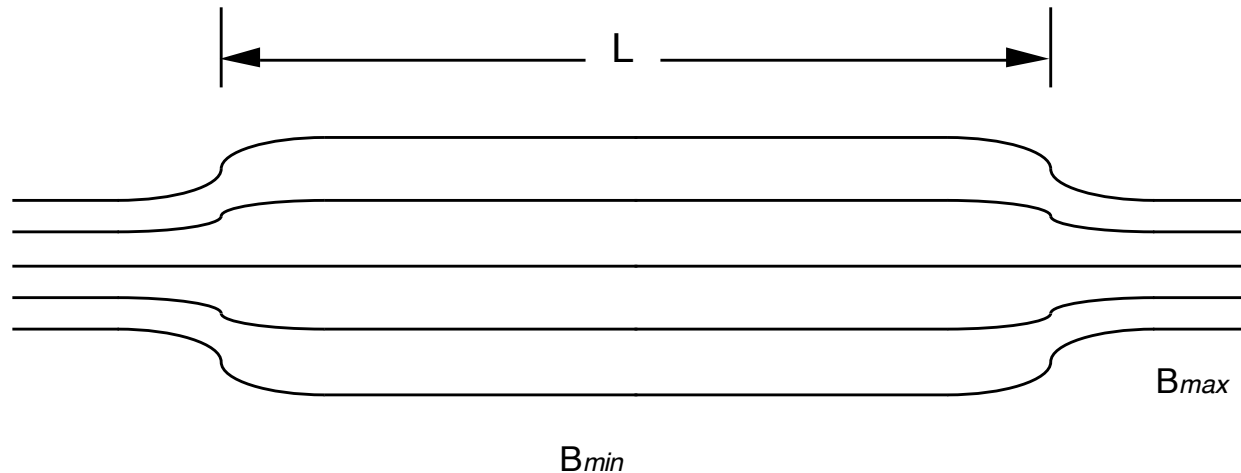
$$\frac{1}{r} \frac{\partial}{\partial r} (r B_r) + \frac{\partial B_z}{\partial z} + \frac{1}{r} \frac{\partial B_\theta}{\partial \theta} = 0 \Leftrightarrow \frac{\partial}{\partial r} (r B_r) = -r \frac{\partial B_z}{\partial z}$$

$$r B_r = -\frac{1}{2} r^2 \frac{\partial B_z}{\partial z} \Leftrightarrow B_r = -\frac{1}{2} r \frac{\partial B_z}{\partial z}$$

$$\vec{F} = q \vec{v} \times \vec{B} \Rightarrow F_z = q v_\theta B_r \Leftrightarrow$$

$$F_z = -\frac{1}{2} q v_\theta r_L \frac{\partial B_z}{\partial z} = -\frac{1}{2} q v_\perp \frac{m v_\perp}{|q| B} \frac{\partial B_z}{\partial z} = -\mu \frac{\partial B_z}{\partial z}$$

# Mirror confinement (mirror-reflected particles)

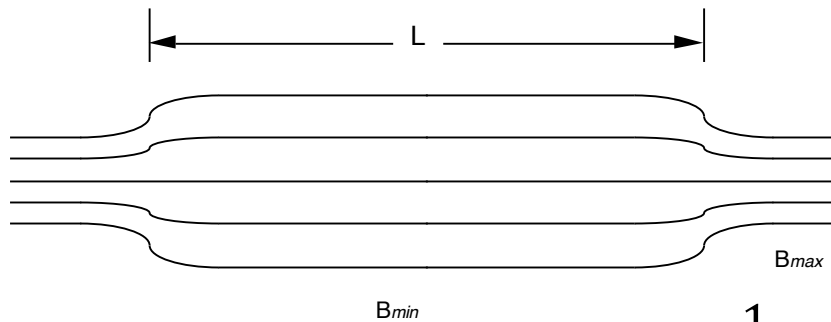


$$\mu B = \frac{1}{2} m v_{\perp}^2, \mu = \text{constant}$$

- Now that we believe in the mirror force, we see that it can be used for confinement.
- We also recognize that  $\mu B$  is the potential energy of the guiding center particle due to its magnetic moment.
- A guiding center particle is in a potential well in a mirror device (above)
- For which particles is the potential well sufficiently deep for trapping?



# Mirror confinement (mirror-reflected particles)



$$\frac{\partial \vec{B}}{\partial t} = 0$$

$$\mu B = \frac{1}{2} m v_{\perp}^2, \mu = \text{constant}$$

$$\frac{1}{2} m v^2 = \frac{1}{2} m v_{\perp}^2 + \frac{1}{2} m v_{\parallel}^2 = \mu B + \frac{1}{2} m v_{\parallel}^2 = \text{constant} = E_k$$

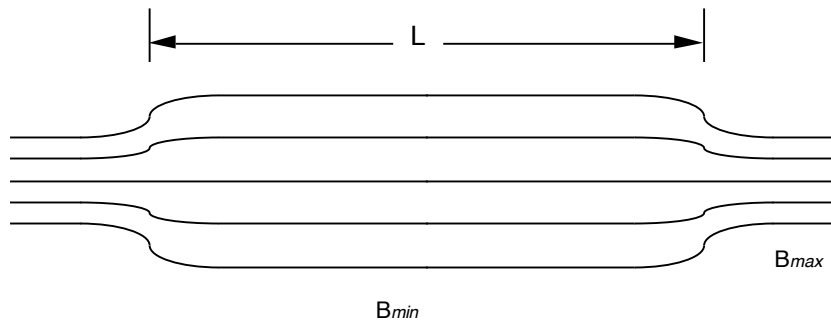
$$\mu B \leq E_k$$

If  $\mu B_{max} > E_k$  then the particle is trapped, otherwise it passes through

$$\mu = \frac{1}{2} m v_{\perp 0}^2 / B_{min}$$

$$\frac{1}{2} \frac{m v_{\perp 0}^2}{B_{min}} B_{max} > \frac{1}{2} m v_0^2 \Leftrightarrow \frac{B_{max}}{B_{min}} > \frac{v_0^2}{v_{\perp 0}^2}$$

# Mirror confinement (mirror-reflected particles)



$$\frac{\partial \vec{B}}{\partial t} = 0$$

$$\frac{B_{max}}{B_{min}} > \frac{v_0^2}{v_{\perp 0}^2} = \frac{1}{\sin^2 \theta} \Leftrightarrow \sin^2 \theta > \frac{B_{min}}{B_{max}}$$

- $\theta$  is the angle between the velocity vector and the B-field vector
- Only part of the Maxwellian is confined – there are always particles that have a small pitch angle
- The mirror ratio  $B_{max}/B_{min}$  is limited by available magnet technology and by needing a minimum (!)  $B_{min}$  so the particles are confined in the long region
- The trapped particles perform a periodic motion in the parallel direction giving rise to a second adiabatic invariant  $J = \oint_0^L v_{||} dz$
- An electrostatic potential can add (or subtract) to the potential well:  $q\phi + \mu B$
- For uniform B (mirror ratio 1) we must use an electrostatic potential: Penning trap

# The parallel Larmor radius

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- For the magnetic moment to be conserved, the gyration needs to be a nearly periodic motion. We earlier required that the Larmor radius be small compared to the distance over which the B-field changes:

$$\frac{B}{|\nabla B|} \gg r_L = \frac{mv_{\perp}}{qB}$$

- We clearly must require also that the particle does not move into a region with a substantially different magnetic field (direction or strength) in a single gyration.

$$\frac{B}{|\nabla B|} \gg \frac{2\pi v_{\parallel}}{\omega_c} = 2\pi \frac{mv_{\parallel}}{qB} = 2\pi r_{L\parallel}$$

- Thus, the “parallel Larmor radius” must also be small compared to the characteristic scale length over which B changes

# Time varying fields

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- One can show that  $\mu$  is conserved also when  $dB/dt$  is nonzero, as long as the B field variation is small in one gyration and has no frequency component near  $\omega_c$
- Implies that perpendicular kinetic energy changes, which is due to E:

$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

- A time varying E-field gives rise to a so-called polarization drift (in addition to an altered ExB drift):

$$\vec{v}_D = \frac{m}{qB^2} \frac{\partial \vec{E}}{\partial t} \text{ if } \omega \ll \omega_c$$

- For  $\omega = \omega_c$  there is no  $\mu$  conservation, and the polarization drift formula is no longer valid:
  - Cyclotron resonance (can be used for plasma heating for example)

# Summary

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- In order to confine charged particles in a magnetic field for many single particle collisions:
  - Must have a small Larmor radius
- In this limit, full orbit calculations are very expensive, and generally not necessary
- The guiding center approximation is very useful here:
  - Charged gyrating particle is approximated as a charged current ring, with constant anti-aligned magnetic dipole moment, sliding along the magnetic field line in the ring center
  - Such particles have slow drifts away from their “birth” magnetic field line, which can be calculated analytically
  - Can calculate complicated trajectories with relative ease
  - Can identify good magnetic traps (particles confined for many collisions) and ones that are less successful...

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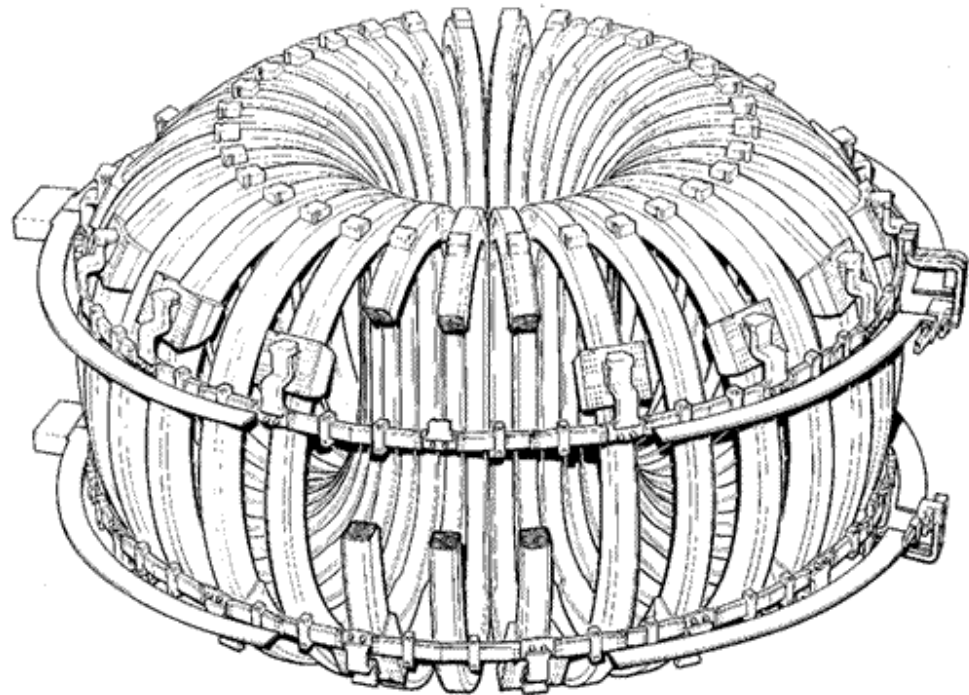
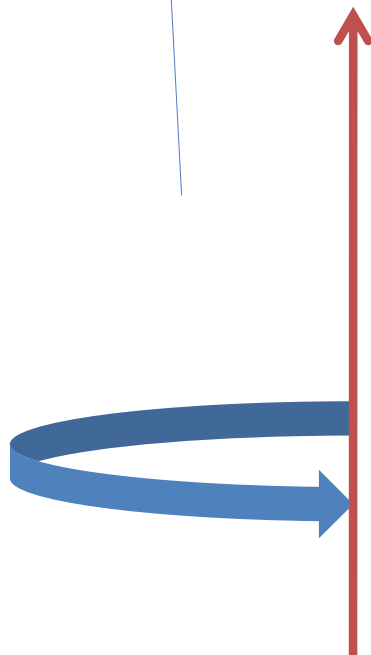
**Examples in the following slides:**

# Pure toroidal field trap (Example: neutral plasma)

- The magnetic field from a dense toroidal set of coils is equivalent to that from an infinite, straight current carrying wire
- Particles experience magnetic drifts in the vertical direction:

$$\vec{v}_{\nabla B + R_C} = \frac{2mv_{\parallel}^2 + mv_{\perp}^2}{2B} \frac{\vec{B} \times \nabla B}{qB^2}$$

Opposite for ions and electrons – they drift apart

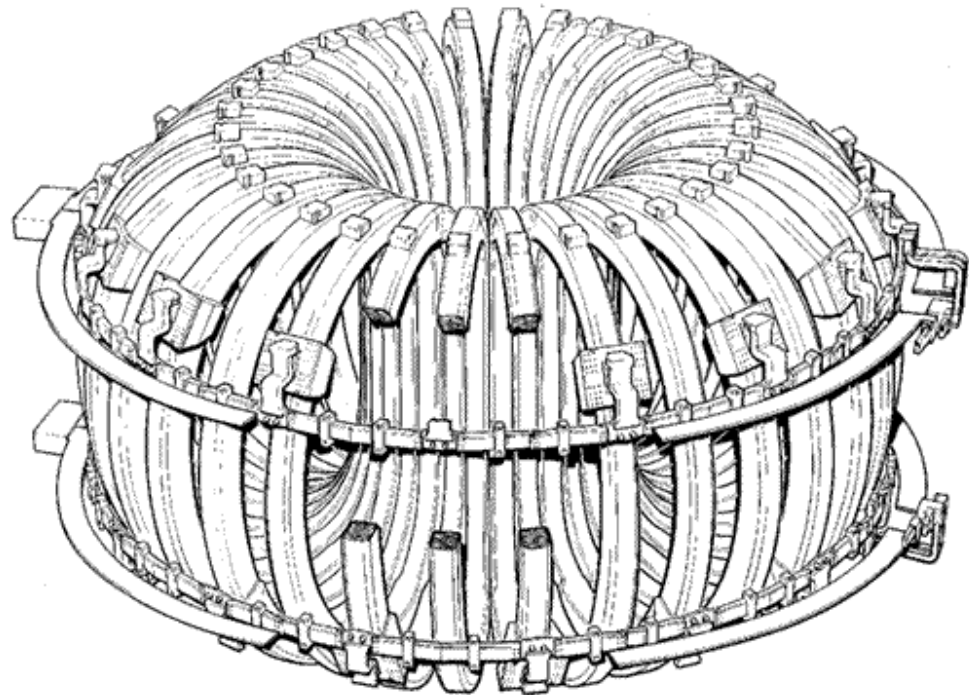
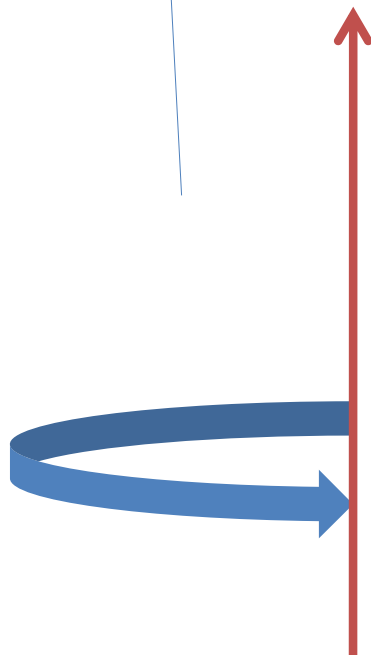


# Pure toroidal field trap (Example: neutral plasma)

- Opposite for ions and electrons – they drift apart

$$\vec{v}_{\nabla B + R_C} = \frac{2mv_{\parallel}^2 + mv_{\perp}^2}{2B} \frac{\vec{B} \times \nabla B}{qB^2}$$

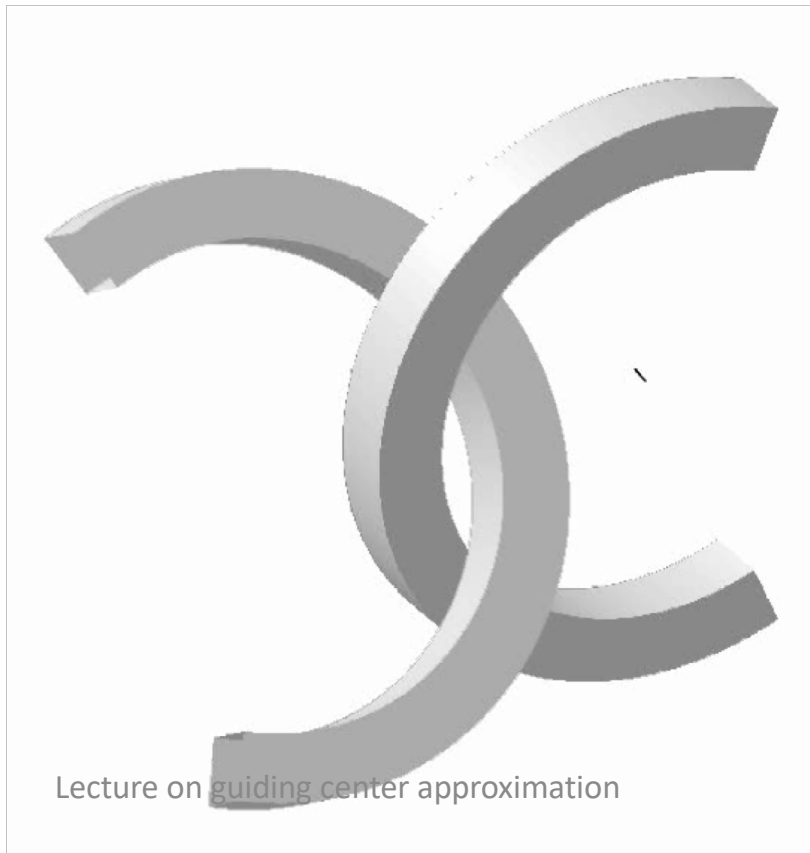
- Vertical electric field sets up:  $\vec{v}_E = \frac{E\hat{z} \times B\hat{\phi}}{B^2} = \frac{E}{B} \hat{R}$
- Outward radial expansion of plasma – no confinement





# Non-neutral plasmas in a stellarator

- A stellarator is a magnetic surface configuration: Each magnetic field line wraps around a toroidal surface, never leaving the surface.
- Also mainly toroidal field – also vertical drift of particles?
- No – vertical drifts cancel because of the poloidal motion that the particle has, as a result of parallel motion along the magnetic field



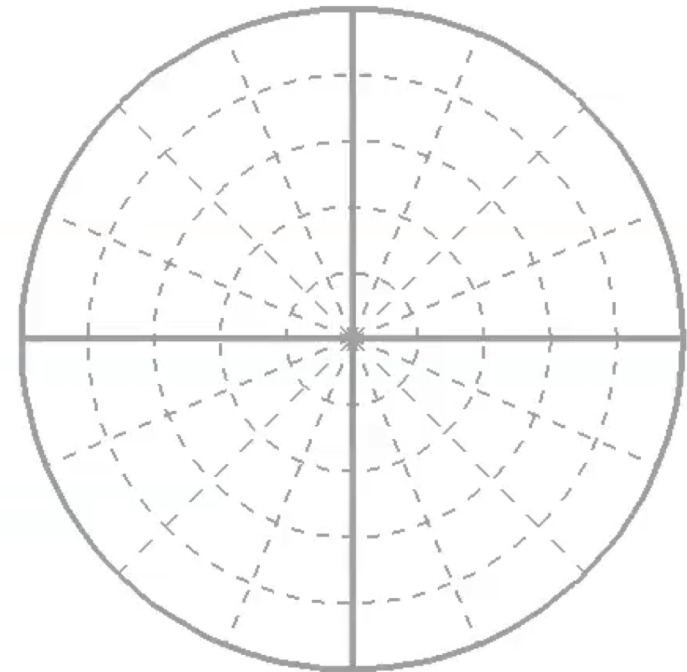
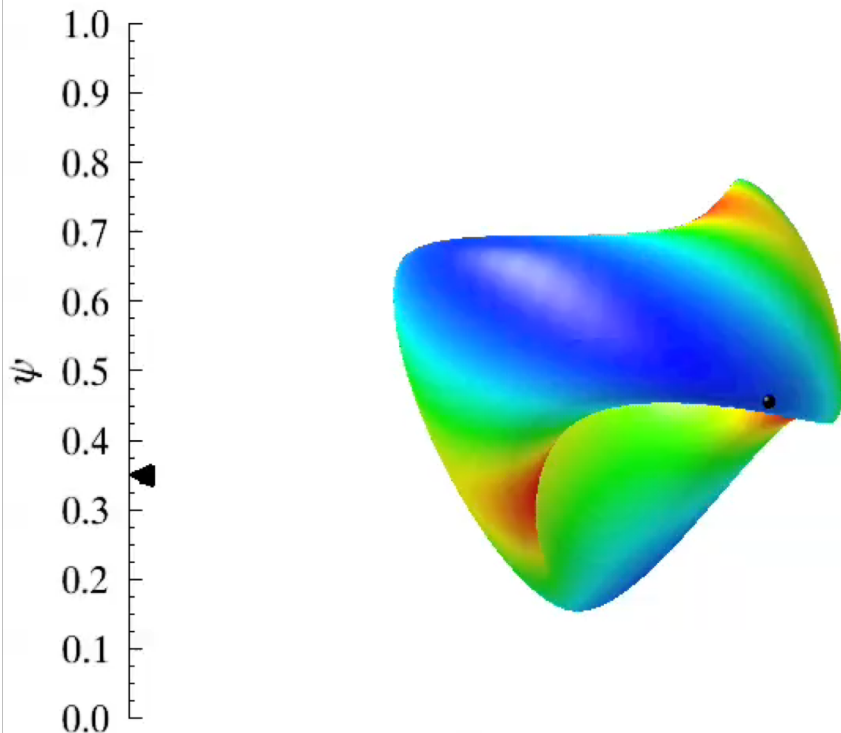
# Without E-field, CNT has “bad” orbits!

CNT is a “classical stellarator” – will not work well for fusion:

About 50% of particles are magnetically trapped (due to mirror force/first adiabatic invariant).

They don’t circulate toroidally, therefore don’t circulate poloidally, and drift out of CNT due to the magnetic drifts. Example:

$t = 0.00\mu\text{s}$

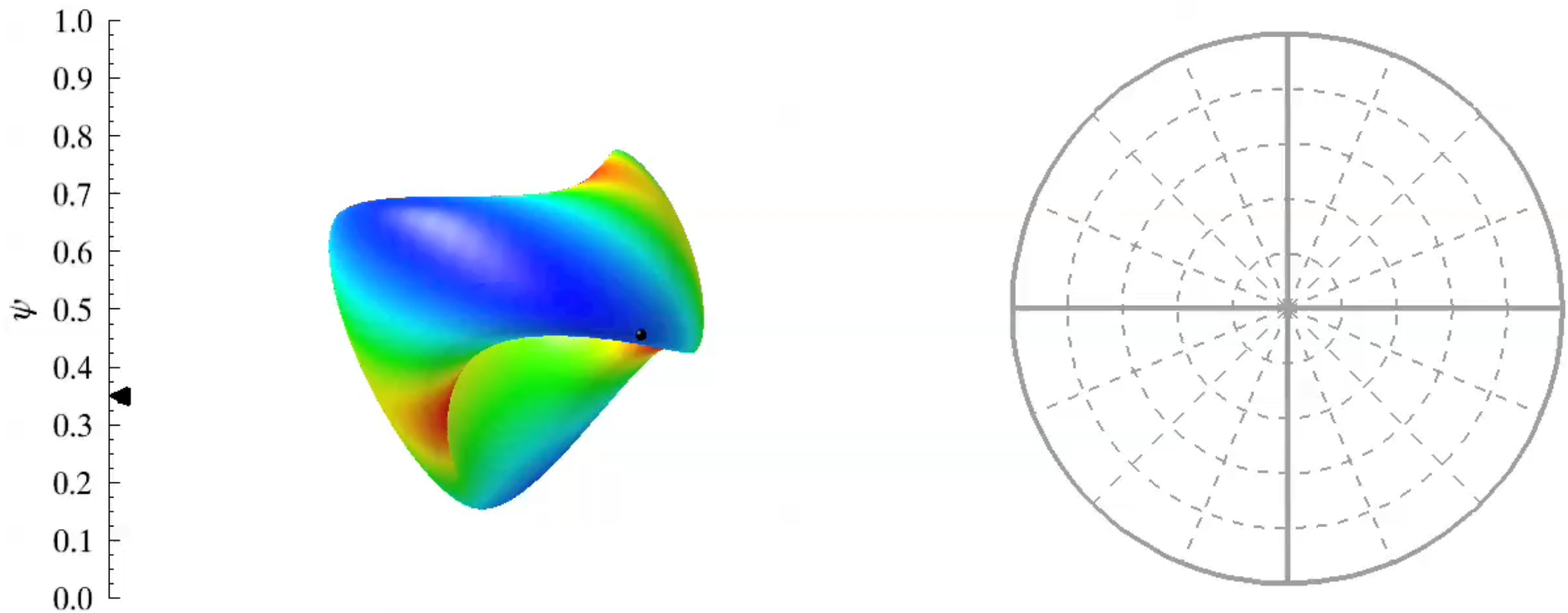


## ExB could come to the rescue

A strong space charge electric field – constant on a magnetic surface – is added to the simulation of the trapped particle

Now it is confined! For much the same reasons as in the pure toroidal field trap

$t = 0.00\mu\text{s}$



# Electric fields in stellarators

How large of a role does the bulk ExB drift play relative to the magnetic drifts?

$$\left| \frac{v_{ExB}}{v_{\nabla B}} \right| \approx \left| \frac{\nabla \phi / B}{(W_k \nabla B / e B^2)} \right| \approx \left| \frac{e \phi}{W_k} \right|$$

Pure-electron plasma: Dominant (factor of 10-1000)

Thermal particles in a quasineutral plasma: Depends.. (0.2-5)

Set by ambipolarity: In steady state, the positive and negative charges of the plasma must leave at the same rate (they “arrive” by neutralization of atoms – ie. at the same rate).

If one species has a tendency to be less well confined (higher mass, higher temperature etc) it will initially leave faster, leaving space charge electric fields that begin to hold it back.

For some plasmas,  $T_e \gg T_i$  which drives a relatively strong positive radial electric field to “hold electrons in and push ions out” – but actually typically improves confinement of both species

For  $T_e \sim T_i$ , usually a negative radial electric field develops

The orbit healing magic of a radial electric field cannot “fix”  $\alpha$ -confinement in a future reactor:

Ratio is negligibly small:  $\sim 35 \text{ keV} / 3.5 \text{ MeV} \sim 0.01$